

Species Tag:	19004	Name:	H <sub>3</sub> O <sup>+</sup>
Version:	3		Hydronium ion,
Date:	Jan. 2010		GS, $\nu_2 = 1$
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Lines Listed:	304	Q(300.0)=	520.1679
Freq. (GHz) <	10000	Q(225.0)=	326.8681
Max. J:	20	Q(150.0)=	174.0742
LOGSTR0=	-5.0	Q(75.00)=	62.5579
LOGSTR1=	-5.0	Q(37.50)=	26.6647
Isotope Corr.:	0.0	Q(18.75)=	15.9161
Egy. (cm <sup>-1</sup> ) >	0.0	Q(9.375)=	11.6139
$\mu_a$ =		A=	331411.77
$\mu_b$ =		B=	A
$\mu_c$ =	1.438	C=	186836.77

The two inversion states of the ground state,  $\nu_1 = 1$ ,  $\nu_2 = 1$ ,  $\nu_3 = 1$ ,  $\nu_4 = 1$  states are included in this calculation. The vibrational designations are as the following: 00 for  $0^-$ ; 01 for  $0^+$ ; 02 for  $\nu_2^-$ , 03 for  $\nu_2^+$ , 04 for  $\nu_1^-$ , 05 for  $\nu_1^+$ , 06 for  $\nu_3^-$ ,  $l = 1$ ; 07 for  $\nu_3^-$ ,  $l = -1$ ; 08 for  $\nu_3^+$ ,  $l = 1$ ; 09 for  $\nu_3^+$ ,  $l = -1$ ; 10 for  $\nu_4^-$ ,  $l = 1$ ; 11 for  $\nu_4^-$ ,  $l = -1$ ; 12 for  $\nu_4^+$ ,  $l = 1$ ; 13 for  $\nu_4^+$ ,  $l = -1$ ; The experimental measurements were reported by Plummer et al., 1985, J. Chem. Phys. **83**, 1428; Bogey et al., 1985, Astron. Astrophys. **148**, L11; Verhoeve et al., 1988, Chem. Phys. Lett., **143**, 501; Verhoeve et al., 1989, Chem. Phys. Lett., **161**, 195; Liu and Oka, 1985, Phys. Rev. Lett. **54**, 1787; Liu et al., 1985, J. Chem. Phys. **82**, 5368; Liu et al., 1986, J. Chem. Phys., **84**, 1312; Davies et al., 1985, J. Opt. Soc. Am. B, **2**, 794; Haese and Oka, 1984, J. Chem. Phys., **80**, 572; Lemoine and Destombes, 1984, Chem. Phys. Lett., **111**, 284; Zheng et al., 2007, Chin. Phys. Lett., **24**, 2569; Tang and Oka, 1999, J. Mol. Spectrosc., **196**, 120; Stahn et al., 1987, Mol. Phys. **60**, 121; Ho al., 1991, J. Mol. Spectrosc., **149**, 530; Begemann MH, Saykally RJ, J. Chem. Phys. **82**(8), 3570-3579, 1985; Gruebele et al., 1987 J. Chem. Phys., **87**, 3347; Stephenson SK, Saykally RJ, Chem. Rev. 105(9), 3220-3234, 2005; Uy D, White ET, Oka T, J. Mol. Spectrosc., 183(2), 240-244, 1997; Yu et al., 2008, Astrophys. J. Supp. 180(1), 119, 2009.

The dipole moments for the inversion transitions,  $\nu_1$  and  $\nu_2$  were taken from the results from Botschwina et al. 1984, Chem. Phys. Lett. **102**, 299; the dipole moments for  $\nu_3$  and  $\nu_4$  are not available and were estimated based on the available infrared intensity ratios (Colvin et al. 1983, J. Chem. Phys. **79**, 1551 ) with the assumption that infrared intensities are proportional to the square of dipole moments.